



1 **Inter-comparison and improvement of 2-stream shortwave radiative transfer**
2 **models for unified treatment of cryospheric surfaces in ESMs**

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11 **Abstract.** Snow is an important climate regulator because it greatly increases the surface
12 albedo of large parts of the Earth. Earth System Models (ESMs) often adopt 2-stream
13 approximations with different radiative transfer techniques, the same snow therefore has
14 different solar radiative properties depending whether it is on land or on sea ice. Here we
15 inter-compare three 2-stream algorithms widely used in snow models, improve their
16 predictions at large zenith angles, and introduce a hybrid model suitable for all
17 cryospheric surfaces in ESMs. The algorithms are those employed by the SNOW ICe and
18 Aerosol Radiative (SNICAR) module used in land models, and by Icepack, the column
19 physics used in the Los Alamos sea ice model CICE and MPAS-seaice, and a 2-stream
20 discrete ordinate (2SD) model. Compared with a 16-stream benchmark model, the errors
21 in snow visible albedo for a direct-incident beam from all three 2-stream models are
22 small ($< \pm 0.005$) and increase as snow shallows, especially for aged snow. The errors in
23 direct near-infrared (near-IR) albedo are small ($< \pm 0.005$) for solar zenith angles $\theta < 75^\circ$,
24 and increase as θ increases. For diffuse incidence under cloudy skies, Icepack produces
25 the most accurate snow albedo for both visible and near-IR ($< \pm 0.0002$) with the lowest
26 underestimate (-0.01) for melting thin snow. SNICAR performs similarly to Icepack for
27 visible albedos, with a slightly larger underestimate (-0.02), while it overestimates the
28 near-IR albedo by an order of magnitude more (up to 0.04). 2SD overestimates both
29 visible and near-IR albedo by up to 0.03. We develop a new parameterization that adjusts
30 the underestimated direct near-IR albedo and overestimated direct near-IR heating
31 persistent across all 2-stream models for solar zenith angles $> 75^\circ$. These results are
32 incorporated in a hybrid model SNICAR-AD, which can now serve as a unified solar
33 radiative transfer model for snow in ESM land, land ice, and sea-ice components.

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35 1. Introduction

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Snow cover on land, land ice, and sea ice, modulates the surface energy balance of large parts of the Earth, principally because even a thin layer of snow greatly increases the surface albedo. Integrated over the solar spectrum, the broadband albedo of opaque snow ranges from 0.7 – 0.9 (e.g., Wiscombe and Warren 1980; Dang et al., 2015). In contrast, the albedo of other natural surfaces is smaller: 0.2, 0.25, and 0.5-0.7 for damp soil, grassland, and bare multi-year sea ice, respectively (Perovich 1996; Liang et al., 2002; Brandt et al., 2005; Bøggild et al., 2010). An accurate simulation of the shortwave radiative properties of snowpack is therefore crucial for spectrally partitioning solar energy and representing snow-albedo feedbacks across the Earth system. Unfortunately, computational demands and coupling architectures often constrain representation of snowpack radiative processes in Earth System Models (ESMs) to relatively crude approximations such as 2-stream methods (Wiscombe and Warren, 1980, Toon et al., 1989). In this work, we inter-compare 2-stream methods widely used in snow models and then introduce a new parameterization that significantly reduces their snowpack reflectance and heating biases at large zenith angles, to produce more realistic behavior in polar regions.

Snow albedo is determined by many factors including the snow grain radius, the solar zenith angle, cloud transmittance, light-absorbing particles, and the albedo of underlying ground if snow is optically thin (Wiscombe and Warren, 1980; Warren and Wiscombe, 1980); it also varies strongly with wavelength since the ice absorption coefficient varies by 7 orders of magnitudes across the solar spectrum (Warren and Brandt, 2008). At visible wavelengths (0.2 - 0.7 μm), ice is almost non-absorptive so that the absorption of visible energy by snowpack is mostly due to the light-absorbing particles (e.g. black carbon, organic carbon, mineral dust) that were incorporated during ice nucleation in clouds, scavenged during precipitation, or slowly sedimented from the atmosphere by gravity (Warren and Wiscombe, 1980, 1985; Doherty et al., 2010, 2014, 2016; Wang et al., 2013; Dang and Hegg 2014). As snow becomes shallower, visible photons are more likely to penetrate through snowpack and get absorbed by darker underlying ground. At near-infrared (near-IR) wavelengths (0.7 – 5 μm), ice is much more absorptive and the snow albedo is lower than the visible albedo. Larger ice crystals form a lower albedo surface than smaller ice crystals hence aged snowpacks absorb more solar energy. Photons incident at smaller solar zenith angles are more likely to penetrate deeper vertically and be scattered in the snowpack until being absorbed by the ice/the



71 underlying ground/absorbing impurities, which also leads to a smaller snow albedo.
72 To compute the reflected solar flux, spectrally resolved albedo must be weighted by the
73 incident solar flux, which is mostly determined by solar zenith angle, cloud cover and
74 transmittance, and column water vapor. Modeling the solar properties of snowpacks must
75 consider the spectral signatures of these atmospheric properties.

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77 Several parameterizations have been developed to compute the snow solar properties
78 without solving the radiative transfer equations and some are incorporated into ESMs or
79 regional models. Marshall and Warren (1987) and Marshall (1989) parameterized snow
80 albedo in both visible and near-IR bands as functions of snow grain size, solar zenith
81 angle, cloud transmittance, snow depth, underlying surface albedo, and black carbon
82 content. Marshall and Oglesby (1994) used this in an ESM. Gardner and Sharp (2010)
83 parameterized the all-wave snow albedo with similar inputs. This was incorporated into
84 the regional climate model RACMO
85 (<https://www.projects.science.uu.nl/iceclimate/models/racmo.php>) to simulate snow
86 albedo in glaciated regions like Antarctica and Greenland (Munneke et al., 2011). Dang
87 et al., (2015) compute snow albedo as functions of snow grain radius, black carbon
88 content, and dust content for visible and near-IR bands and 14 narrower bands used in the
89 rapid radiative transfer model (RRTM, Mlawer and Clough, 1997). Their
90 parameterization can also be expanded to different solar zenith angles using the zenith
91 angle parameterization developed by Marshall and Warren (1987). Aoki et al., (2011)
92 developed a more complex model (PBSAM) based on the offline snow albedo and a
93 transmittance look-up table. This can be applied to multilayer snowpack to compute the
94 snow albedo and the solar heating profiles as functions of snow grain size, black carbon
95 and dust content, snow temperature, and snowmelt water equivalent. These
96 parameterizations are often in the form of simplified polynomial equations, and are
97 especially suitable to long-term ESM simulations that require less time-consuming snow
98 representations.

99
100 More complex models that explicitly solve the multiple scattering radiative transfer
101 equations have also been developed to compute snow solar properties. Flanner and
102 Zender (2005) developed the SNow Ice and Aerosol Radiation model (SNICAR) that
103 utilizes 2-stream approximations (Wiscombe and Warren 1980; Toon et al., 1989) to
104 predict heating and reflectance for multi-layer snowpack. They implemented SNICAR in
105 the Community Land Model (CLM) to predict snow albedo and vertically-resolved solar
106 absorption for snow-covered surfaces. Before SNICAR, CLM prescribed snow albedo



107 and confined all solar absorption to the top snow layer (Flanner and Zender 2005). Over
108 the past decades, updates and new features have been added to SNICAR to consider more
109 processes such as black carbon/ice mixing states (Flanner et al., 2012) and snow grain
110 shape (He et al., 2018b). Concurrent with the development of SNICAR, Briegleb and
111 Light (2007) improved the treatment of sea-ice solar radiative calculations in Community
112 Climate System Model (CCSM). They implemented a 2-stream delta-Eddington method
113 that allows CCSM to compute bare/ponded/snow-covered sea ice albedo and solar
114 absorption profiles of multi-layer sea ice. Before these improvements, the sea-ice albedo
115 was computed based on surface temperature, snow thickness, and sea-ice thickness using
116 averaged sea ice and snow albedo. This method has carried into the sea-ice physics
117 library Icepack (<https://github.com/CICE-Consortium/Icepack/wiki>) that comprises the
118 column physics used by the Los Alamos Sea Ice Model CICE (Hunke et al., 2010) and
119 MPAS-seaice (Turner et al., 2018). CICE itself is used in numerous global and regional
120 models.

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122 The shortwave methods in SNICAR and in CICE solve the multiple scattering radiative
123 transfer equations and provide much improved solar radiative representations for the
124 cryosphere, though their separate development and implementation created an artificial
125 divide for snow simulation. In ESMs that utilize both SNICAR and CICE/MPAS-seaice,
126 such as the Community Earth System Model (CESM, <http://www.cesm.ucar.edu/>) and the
127 Energy Exscale Earth System Model (E3SM, previously known as ACME,
128 <https://e3sm.org/>), the solar radiative properties of snow on land and snow on sea ice are
129 computed separately via SNICAR and CICE/MPAS-seaice. As a result, the same snow in
130 nature has different solar radiative properties such as reflectance depending on which
131 model represents it. These differences are model artifacts that should be eliminated so
132 that snow has consistent properties across the Earth system.

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134 In this paper, we evaluate the accuracy and biases of three 2-stream algorithms described
135 in Section 2 and Table 1, including the algorithms used in SNICAR and Icepack, at
136 representing reflectance and heating. We use these results to develop and justify a unified
137 surface shortwave radiative transfer method for all Earth system model components in the
138 cryosphere.

139

140 **2. Radiative Transfer Model**

141

142 In this section, we summarize the three 2-stream models and the benchmark DISORT



143 model with 16-streams. These algorithms are well documented in papers by Toon et al.,
 144 (1989), Briegleb and Light (2007), Jin and Stamnes (1994), and Stamnes et al. (1988).
 145 Readers interested in detailed mathematical derivations should refer to those papers. We
 146 only include their key equations to illustrate the difference among 2-stream models for
 147 discussion purposes.

148 2.1 SNICAR

150 SNICAR adopts the 2-stream algorithms and the rapid solver developed by Toon et al.,
 151 (1989) to compute the solar properties of multi-layer snowpacks. These 2-stream
 152 algorithms are derived from the general equation of radiative transfer in a plane parallel
 153 media:

$$155 \mu \frac{\partial I}{\partial \tau}(\tau, \mu, \Phi) = I(\tau, \mu, \Phi) - \frac{\omega}{4\pi} \int_0^{2\pi} \int_{-1}^1 P(\mu, \mu', \phi, \phi') I(\tau, \mu', \Phi') d\mu' d\phi' - S(\tau, \mu, \Phi)$$

156
157 (1)
158
159

160 where $\arccos(\mu)$ and Φ are zenith angle and azimuth angle, ω is single-scattering albedo.
 161 On the right-hand side, the three terms are intensity at optical depth τ , internal source
 162 term due to multiple scattering, and external source term S . For a purely external source
 163 at solar wavelengths S is:

$$165 S = \frac{\omega}{4} F_s P(\mu, -\mu_0, \phi, \phi_0) \exp\left(\frac{-\tau}{\mu_0}\right)$$

166 (2)

167 where πF_s is incident solar flux, μ_0 is the incident direction of the solar beam. Integrating
 168 equation (1) over azimuth and zenith angles yields the general solution of 2-stream
 169 approximations (Meador and Weaver, 1980). The upward and downward fluxes at optical
 170 depth τ of layer n can be represented as:

$$173 F_n^+ = k_{1n} \exp(\Lambda_n \tau) + \Gamma_n k_{2n} \exp(-\Lambda_n \tau) + C_n^+(\tau)$$

174 (3a)

$$175 F_n^- = \Gamma_n k_{1n} \exp(\Lambda_n \tau) + k_{2n} \exp(-\Lambda_n \tau) + C_n^-(\tau)$$

176 (3b)
177

178 where Λ_n , Γ_n , C_n are known coefficients determined by the 2-stream method, incident
 179 solar flux, and solar zenith angle; whereas k_{1n} and k_{2n} are unknown coefficients



180 determined by the boundary conditions. For an N-layer snowpack, the solutions for
 181 upward and downward fluxes are coupled at layer interfaces to generate 2N equations
 182 with 2N unknown coefficients k_{1n} and k_{2n} . Combining these equations linearly generates
 183 a new set of equations with terms in tridiagonal form that enables the application of a fast
 184 tri-diagonal matrix solver. With the solved coefficients, the upward and downward fluxes
 185 are computed at different optical depths (Equations 3a and 3b) and eventually the
 186 reflectance, transmittance, and absorption profiles of solar flux for any multilayer
 187 snowpack.

188

189 SNICAR itself implements all three 2-stream algorithms in Toon et al., (1989):
 190 Eddington, Quadrature, and Hemispheric-mean. In ESM simulations, it utilizes the
 191 Eddington and Hemispheric-mean approximations to compute the visible and near-IR
 192 snow properties, respectively (Flanner et al., 2007). In addition to their algorithms,
 193 SNICAR implements a Delta-transform of the fundamental input variables asymmetry
 194 factor (g), single-scattering albedo (ϖ), and optical depth (τ) to account for the strong
 195 forward scattering in snow (Equations 2 (a)-(c), Wiscombe and Warren, 1980).

196

197 2.2. Icepak, CICE, and MPAS-seaice

198 Icepak, CICE and MPAS-seaice use the same solar radiative treatment developed and
 199 documented by Briegleb and Light (2007). In the following discussions, we will refer to
 200 this method as CICE since it is more widely used. Sea ice is divided into multiple layers
 201 to first compute the single-layer reflectance and transmittance using 2-stream delta-
 202 Eddington solutions to account for the multiple scattering of light within each layer
 203 (Equation set 50, Briegleb and Light, 2007), where the name “delta” implies CICE
 204 implements the Delta-transform to account for the strong forward scattering of snow and
 205 sea ice (Equations 2 (a)-(c), Wiscombe and Warren, 1980). The direct albedo and
 206 transmittance are computed by equations:

207

$$208 \quad R(\mu_{0,n}) = A_n \exp\left(\frac{-\tau}{\mu_{0,n}}\right) + B_n (\exp(\varepsilon_n \tau) - \exp(-\varepsilon_n \tau)) - K_n \quad (4a)$$

209

$$210 \quad T(\mu_{0,n}) = E_n + H_n (\exp(\varepsilon_n \tau) - \exp(-\varepsilon_n \tau)) \exp\left(\frac{-\tau}{\mu_{0,n}}\right) \quad (4b)$$

211

212 where coefficients A_n , B_n , K_n , E_n , H_n , and ε_n are determined by the single-scattering
 213 albedo (ϖ), asymmetry factor (g), optical depth (τ), and angle of incident beam at layer n
 214 ($\mu_{0,n}$). Following the delta-Eddington assumption, simple formulas are available for the



215 single-layer reflectance and transmittance under both clear sky (direct flux, equations 4a
216 and 4b) and overcast sky (diffuse flux) conditions, however, the formula derived by
217 applying diffuse-flux upper boundary conditions sometimes yields negative albedos
218 (Wiscombe 1977). To avoid the unphysical values, diffuse reflectance \bar{R} and
219 transmittance \bar{T} of a single layer are computed by integrating the direct reflectance $R(\mu)$
220 and transmittance $T(\mu)$ over the incident hemisphere assuming isotropic incidence:
221

$$222 \quad \bar{R} = 2 \int_0^1 \mu R(\mu) d\mu \quad (5a)$$

$$223 \quad \bar{T} = 2 \int_0^1 \mu T(\mu) d\mu \quad (5b)$$

224
225
226 This is the same as the method proposed by Wiscombe and Warren (1980, their equation
227 5). In practice, eight Gaussian angles are implemented to perform the integration for
228 every layer.

229
230 These layer reflectance and transmittance of direct and diffuse components are then
231 combined to account for the inter-layer scattering of light to compute the reflectance and
232 transmission at every interface (Equation set 51, Briegleb and Light, 2007), and
233 eventually the upward and downward fluxes (Equation set 52, Briegleb and Light, 2007).
234 These upward and downward fluxes at each optical depth are then used to compute the
235 column reflectance and transmittance, and the absorption profiles for any multilayered
236 media, such as snowpacks on land and sea ice.

237
238 In nature, a large fraction of sea ice is covered by snow during winter. As snow melts
239 away in late spring and summer, it exposes bare ice, and melt ponds form on the ice
240 surface. Such variation of sea-ice surface types requires the shortwave radiative transfer
241 model to be flexible and capable of capturing the light refraction and reflection.
242 Refractive boundaries exist where air (refractive index $m_{re} = 1.0$), snow (assuming snow
243 as medium of air containing a collection of ice particles, $m_{re} = 1.0$), pond (assuming pure
244 water, $m_{re} = 1.33$), and ice (assuming pure ice, $m_{re} = 1.31$) are present in the same sea-ice
245 column. The general solution of delta-Eddington, and the 2-stream algorithms used in
246 SNICAR are not applicable to such non-uniformly refractive layered media. To include
247 the effects of refraction, Briegleb and Light (2007) modified the adding formula at the
248 refractive boundaries (i.e. interfaces between air/ice, snow/ice, air/pond). The reflectance
249 and transmittance of the adjacent layers above and below the refractive boundary are



250 combined with modifications to include the Fresnel reflection and refraction of direct and
251 diffuse fluxes (Section 4.1, Briegleb and Light, 2007). This adding-doubling delta-
252 Eddington method can thus be applied to any layered media with either uniform (e.g.,
253 snow on land) or non-uniform (e.g., snow on sea ice) refractive indexes.

254

255 In this paper, we focus on snowpacks that can be treated as uniform refractive media such
256 as the air/snowpack/land columns assumed in SNICAR. An ideal radiative treatment for
257 snow should however keep the potential to include refraction for further applications to
258 snow on sea ice or ice sheets. Therefore, besides these two widely used algorithms in
259 Icepack and SNICAR, we evaluate a third algorithm (section 2.3) that can be applied to
260 layered media with either uniform or non-uniform refractive indexes.

261

262 2.3. 2-stream discrete-ordinate algorithm (2SD)

263 A refractive boundary also exists between the atmosphere and the ocean, and models
264 have been developed to solve the radiative transfer problems in the atmosphere-ocean
265 system using the discrete-ordinate technique (e.g. Jin and Stamnes, 1994; Lee and Liou,
266 2007). Similar to the 2-stream algorithms of Toon et al., (1989) used in SNICAR, Jin and
267 Stamnes (1994) also developed their algorithm from the general equation:

268

$$269 \mu \frac{\partial I}{\partial \tau}(\tau, \mu) = I(\tau, \mu) - \frac{\omega}{4\pi} \int_{-1}^1 P(\tau, \mu, \mu') I(\tau, \mu') d\mu' - S(\tau, \mu) \quad (6)$$

270

271 Equation (6) is the azimuthally integrated version of equation (1). However, for vertically
272 inhomogeneous media like the atmosphere-ocean or sea ice, the external source term
273 $S(\tau, \mu)$ is different. Specifically, for the medium of total optical depth τ^a above the
274 refractive interface, one must consider the contribution from the upward beam reflected
275 at the refractive boundary (second term on the right-hand side):

276

$$277 S^a(\tau, \mu) = \frac{\omega}{4\pi} F_s P(\tau, -\mu_0, \mu) \exp\left(\frac{-\tau}{\mu_0}\right) + \frac{\omega}{4\pi} F_s R(-\mu_0, m) P(\tau, +\mu_0, \mu) \exp\left(\frac{-(2\tau^a - \tau)}{\mu_0}\right)$$

278

279

280

(7)

281 where $R(-\mu_0, m)$ is the Fresnel reflectance of radiation and m is the ratio of the
282 refractive indices of the lower to the upper medium. For the medium below the refractive
283 interface, one must account for the Fresnel transmittance $T(-\mu_0, m)$ and modify the
284 angle of beam travel in media b:



285

$$286 \quad S^b(\tau, \mu) = \frac{\omega}{4\pi} \frac{\mu_0}{\mu_{0n}} F_s T(-\mu_0, m) P(\tau, -\mu_0, \mu) \exp\left(\frac{-\tau^a}{\mu_0}\right) \exp\left(\frac{-(\tau-\tau^a)}{\mu_{0n}}\right) \quad (8)$$

287

288 where μ_{0n} is the cosine zenith angle of refracted beam incident at angle μ_0 above
289 refractive boundary, by Snell's law:

290

$$291 \quad \mu_{0n} = \sqrt{1 - (1 - \mu_0^2)/m^2} \quad (9)$$

292

293 For uniformly refractive media like snow on land, one can just set the refractive index m_{re}
294 equal to 1 for every layer. In this case, the Fresnel reflectance $R(-\mu_0, m)$ is 0 in equation
295 (7), the Fresnel transmittance $T(-\mu_0, m)$ is 1 in equation (8), and μ_{0n} equals to μ_0 : the
296 two source terms $S^a(\tau, \mu)$ and $S^b(\tau, \mu)$ become the same and equal to the source term of
297 homogenous media given in equation (2).

298

299 For 2-stream approximations of this method, analytical solutions of upward and
300 downward fluxes are coupled at each layer interface to generate 2N equations with 2N
301 unknown coefficients for any N-layer stratified column. The solutions of 2-stream
302 algorithms and boundary conditions for homogenous media are well documented
303 (Sections 8.4 and 8.10 of Thomas and Stamnes, 1999). Despite the extra source terms,
304 these 2N equations can also be organized into a tridiagonal matrix similar to the method
305 of Toon et al. (1989) used in SNICAR. Flexibility and speed therefore make this 2-stream
306 discrete-ordinate algorithm (hereafter, 2SD) a potentially good candidate for long-term
307 Earth system modeling. In this work, we only apply 2SD to snowpack and note that it can
308 be applied to any uniformly or non-uniformly refractive media like snow on land or sea
309 ice, with the Delta-transform implemented to fundamental optical variables (Equations 2
310 (a)-(c), Wiscombe and Warren, 1980).

311

312 2.4 16-stream DISORT

313 Besides the mathematical technique, the accuracy and speed of radiative transfer
314 algorithms depend on the number of angles used for flux estimation in the upward and
315 downward hemispheres. The algorithms used in SNICAR, Icepak, and 2SD use one
316 angle to represent upward flux and one angle to represent downward flux, hence they are
317 named 2-stream algorithm. Lee and Liou (2007) use two upward and two downward
318 streams. Jin and Stamnes (1994) documented the solutions for any even number of



319 streams. The speed of these models is slower than 2-stream models while their accuracy
320 is better. To quantify the accuracy of the three 2-stream algorithms for snow shortwave
321 simulations, we use the 16-stream DIScrete-Odinate Radiative Transfer model (DISORT)
322 as the benchmark model (<http://llab.phy.stevens.edu/disort/>) (Stamnes et al., 1988).

323

324 **3. Input for radiative transfer models**

325 In this work, we focus on the performance of 2-stream algorithms for pure snow
326 simulations. The inputs for these three models are the same: single-scattering properties
327 (SSPs, i.e. single-scattering albedo ω , asymmetry factor g , extinction coefficient σ_{ext}) of
328 snow determined by snow grain radius r , snow depth, solar zenith angle θ , solar incident
329 flux, and the albedo of underlying ground (assuming Lambertian reflectance of 0.25 for
330 all wavelengths). A Delta-transform is applied to fundamental input optical variables for
331 all simulations (Equations 2 (a)-(c), Wiscombe and Warren, 1980).

332

333 In snow, photon scattering occurs at the air-ice interface, and the absorption of photons
334 occurs within the ice crystal. The most important factor that determines snow shortwave
335 properties is the ratio of total surface area to total mass of snow grains, aka “the specific
336 surface area” (e.g. Matzl and Schneebeli, 2006, 2010). The specific surface area (β) can
337 be converted to a radiatively effective snow grain radius r :

338

$$339 \quad \beta = 3 / (r \rho_{ice}) \quad (10)$$

340

341 where ρ_{ice} is the density of pure ice, 917 kg m^{-3} . Assuming the grains are spherical, the
342 SSPs of snow can thus be computed using Mie theory (Wiscombe, 1980) and ice optical
343 constants (Warren and Brandt, 2008). In nature, snow grains are not spherical, and many
344 studies have been carried out to quantify the accuracy of such spherical representations
345 (Grenfell and Warren, 1999; Neshyba et al., 2003; Grenfell et al., 2005). In recent years,
346 more research has been done to evaluate the impact of grain shape on snow shortwave
347 properties (Dang et al., 2016; He et al., 2017, 2018ab), and they show that non-spherical
348 snow grain shapes mainly alter the asymmetry factor. Dang et al., (2016) also point out
349 that the solar properties of a snowpack consisting of non-spherical ice grains can be
350 mimicked by a snowpack consisting of spherical grains with a smaller grain size by
351 factors up to 2.4. In this work, we still assume the snow grains are spherical, and this
352 assumption does not qualitatively alter our evaluation of the radiative transfer algorithms.

353



354 The input SSPs of snow grains are computed using Mie theory at fine spectral resolution
355 for a wide range of ice effective radius r from 10 to 3000 μm that covers the possible
356 range of grain radius for snow on Earth (Flanner et al., 2007). The same spectral SSPs
357 were also used to derive the band-averaged SSPs of snow used in SNICAR. Note
358 Briegleb and Light (2007) refer to SSPs as inherent optical properties.

359

360 4. Solar spectra used for the spectral integrations

361 In climate modeling, snow albedo computation at fine spectral resolution is expensive
362 and unnecessary. Instead of computing spectrally resolved snow albedo as shown in
363 Figure 1, wider-band solar properties are more practical. For example, CESM and E3SM
364 aggregate the narrow RRTMG bands used for the atmospheric radiative transfer
365 simulation into visible (0.2 - 0.7 μm) and near-IR (0.7 - 5 μm) bands. The land model and
366 sea-ice model thus receive visible and near-IR fluxes as the upper boundary condition,
367 and return the corresponding visible and near-IR albedos to atmosphere model. In
368 practice, these bands are also partitioned into direct and diffuse components. Therefore, a
369 practical 2-stream algorithm should be able to simulate the direct visible, diffuse visible,
370 direct near-IR and diffuse near-IR albedos and absorptions of snow accurately.

371

372 The band albedo α is an irradiance-weighted average of the spectral albedo $\alpha(\lambda)$:

373

$$374 \alpha = \frac{\int_{\lambda_1}^{\lambda_2} \alpha(\lambda) F(\lambda) d\lambda}{\int_{\lambda_1}^{\lambda_2} F(\lambda) d\lambda} \quad (11)$$

375

376

377 In this work, we use the spectral irradiance $F(\lambda)$ generated by the atmospheric DISORT-
378 based Shortwave Narrowband Model (SWNB2) (Zender et al., 1997; Zender, 1999) for
379 typical clear-sky and cloudy-sky conditions of mid-latitude winter as shown in Figure
380 1(a). The total clear-sky down-welling surface flux at different solar zenith angles are
381 also given in Figure 1(b).

382

383

384 5. Model Evaluation

385 5.1 Spectral albedo and reflected solar flux

386 The spectral reflectance of pure deep snow computed using 2-stream models and 16-
387 stream DISORT are shown in Figure 2. The snow grain radius is 100 μm - a typical grain
388 size for fresh new snow. For clear sky with direct beam source (left column), all three 2-
389 stream models show good accuracy at visible wavelengths (0.3 – 0.7 μm), and within this



390 band, the snow albedo is large and close to 1. As wavelength increases, the albedo
391 diminishes in the near-IR band. 2-stream models overestimate snow albedo at these
392 wavelengths, with maximum biases of 0.013 (SNICAR and CICE) and 0.023 (2SD)
393 within wavelength 1 - 1.7 μm . For cloudy-sky cases with diffuse upper boundary
394 conditions, CICE reproduces the snow albedo at all wavelengths with the smallest
395 absolute error (< 0.005), SNICAR and 2SD both overestimate the snow albedo with
396 maximum biases > 0.04 between 1.1-1.4 μm .

397

398 In both sky conditions, the errors of snow albedo are larger at near-IR wavelengths
399 ranging from 1.0-1.7 μm , while the solar incident flux peaks at 0.5 μm then decreases as
400 wavelength increases. The largest error in reflected flux is within the 0.7-1.5 μm band for
401 SNICAR and 2SD, as shown in the 3rd row of Figure 2. CICE overestimate the direct
402 snow albedo mostly at wavelengths larger than 1.5 μm where the error in reflected flux is
403 almost negligible.

404

405 5.2 Broadband albedo and reflected solar flux

406 Integrated over the visible and near-IR wavelengths, the error in band albedos computed
407 using 2-stream models for different cases are shown in Figure 3-6.

408

409 Figure 3 shows the error in direct band albedo for fixed snow grain radius of 100 μm with
410 different snow depth and solar zenith angles. As introduced in Section 2, SNICAR and
411 CICE both use delta-Eddington method to compute the visible albedo. They overestimate
412 the visible albedo for solar zenith angles smaller than 50° by up to 0.005, and
413 underestimate it for solar zenith angles larger than 50° by up to -0.01. 2SD produces
414 similar results for the visible band but at a larger solar zenith angle threshold of 75° . In
415 the near-IR band, SNICAR and 2SD overestimate the snow albedo for solar zenith angles
416 smaller than 70° , beyond this, the error in albedo increases by up to -0.1 as solar zenith
417 angle increases. CICE produces a similar error pattern with a smaller solar zenith angle
418 threshold at 60° . As snow ages, its average grain size increases. For typical old melting
419 snow of grain radius 1000 μm (Figure 4), 2-stream models produce similar errors of
420 direct albedo in all bands. For snow consisting of smaller grain size, 2-stream models
421 produce larger errors for visible albedo. Integrating over the entire solar band, the three 2-
422 stream models evaluated show similar error patterns for direct albedo.

423

424 For a fixed solar zenith angle of 60° , the error of direct albedo for different snow depth
425 and snow grain radii are shown in Figure 5. SNICAR and CICE underestimate the visible



426 albedo in most scenarios, while 2SD overestimates the visible albedo for a larger range of
427 grain radius and snow depth. All three 2-stream models tend to overestimate the near-IR
428 albedo except for shallow snow with large grain radius; the error of 2SD is one order of
429 magnitude larger than that of SNICAR and CICE.

430

431 Figure 6 is similar to Figure 5, but shows the diffuse snow albedo. In the visible band,
432 SNICAR and CICE generate similar errors in that they both underestimate the albedo as
433 snow grain size increases and snow depth decreases. 2SD overestimates the albedo with
434 maximum error of around 0.015. In the near-IR, 2-stream models tend to overestimate
435 snow albedo, while the magnitude of biases produced by SNICAR and 2SD are one order
436 larger than that of CICE with the maximum error of 0.035 generated by SNICAR. As a
437 result, the all-wave diffuse albedos computed using CICE are more accurate than those
438 computed using SNICAR and 2SD.

439

440 Figures 7, 8 and 9 show the errors in reflected shortwave flux caused by snow albedo
441 errors seen in Figures 3, 4, and 6. In general, 2-stream models produce larger errors in
442 reflected direct near-IR flux (Figure 7 and 8), especially with the 2SD model: the
443 maximum overestimate of reflected near-IR flux is 6-8 Wm^{-2} for deep melting snow with
444 solar zenith angle $< 30^\circ$. Errors in reflected direct visible flux are smaller (mostly within
445 $\pm 1 \text{ Wm}^{-2}$) for all models in most scenarios, and become larger (mostly within $\pm 3 \text{ Wm}^{-2}$) as
446 snow grain size increases to 1000 μm if computed using 2SD. As shown in Figure 9, for
447 diffuse flux with solar zenith angle of 60° at TOA, SNICAR and CICE generate small
448 errors in reflected visible flux (mostly within $\pm 1 \text{ Wm}^{-2}$), while 2SD always overestimates
449 reflected visible flux by up to 5 Wm^{-2} . In the near-IR, SNICAR and 2SD overestimate
450 reflected flux by as much as $10\text{-}12 \text{ Wm}^{-2}$; the error in reflected near-IR flux produced by
451 CICE is much smaller, mostly within $\pm 1 \text{ Wm}^{-2}$.

452

453 In general, CICE produces the most accurate albedo and thus reflected flux for both
454 direct and diffuse components. SNICAR is similar to CICE for its accuracy of direct
455 albedo and flux, yet generates large error for diffuse component. 2SD tends to
456 overestimate snow albedo and reflected flux in both direct and diffuse components and
457 shows the largest errors among three 2-stream models. Note that the final errors of snow
458 albedo and reflected solar flux are the weighted sum of direct and diffuse components,
459 and their weights are largely determined by cloud cover fraction (e.g. Figure 6, Dang et
460 al., 2017), which we do not address explicitly in this paper.



461

462 5.3 Band absorption of solar flux

463 Figure 10 shows absorption profiles of shortwave flux computed using the 16-stream
464 DISORT model, with errors in absorbed fractional solar flux computed using 2-stream
465 models. The snowpack is 10-cm deep, and is divided into 5 layers, each 2-cm thick. The
466 snow grain radius is set to 100 μm . The figure shows fractional absorption for snow
467 layers 1-4 and the underlying ground with albedo of 0.25.

468

469 As shown in the first column of Figure 10, for new snow with radius of 100 μm , most
470 solar absorption occurs in the top 2-cm snow layer, where roughly 10% and 15% of
471 diffuse and direct near-IR flux are absorbed and dominate the solar absorption within
472 snowpack. In the second layer (2-4 cm), the absorption of solar flux is less than 1% and
473 gradually decreases within the interior layers. The underlying ground absorbs roughly 2%
474 of solar flux, mostly visible flux that penetrates the snowpack more efficiently. As snow
475 ages and snow grain grows, photons penetrate deeper into the snowpack. For typical old
476 melting snow with radius of 1000 μm , most solar absorption still occurs in the top 2-cm
477 snow layer, where roughly 20% and 14% of diffuse and direct near-IR flux are absorbed.
478 The second snow layer (2-4 cm) absorbs more near-IR solar flux by roughly 2%. More
479 photons are able to penetrate through the snowpack, and results in a high fractionally
480 absorption by the underlying ground, especially for visible band. As snow depth increase,
481 the ground absorption will decrease for both snow radii.

482

483 Comparing to 16-stream DISORT, 2-stream models underestimate (overestimate) the
484 column solar absorptions for new (old) snow, especially for the surface snow layer and
485 ground layer. Overall, CICE gives the most accurate absorption profiles among three 2-
486 stream models, especially for new snow.

487

488 6. Correction for direct albedo for large solar zenith angles

489

490 It has been pointed out in previous studies that the 2-stream approximations become poor
491 as solar zenith angle approaches 90° (e.g. Wiscombe 1997, Warren 1982). As shown in
492 Figures 3 and 4, all three 2-stream models underestimate the direct snow albedo for large
493 solar zenith angles. In the visible band, when snow grain size is small, the error in direct
494 albedo is almost negligible (Figure 3); while as snow ages and snow grains become
495 larger, the error increases yet still remains low if the snow is deep (Figure 4). In the near-
496 IR, the biases of albedo are also larger for larger snow grain radii. For a given snow size,



497 the magnitudes of such biases are almost independent of snow depth, and mainly
 498 determined by the solar zenith angle. In general, the errors of all-wave direct albedo are
 499 mostly contributed by the errors of near-IR albedo, especially for optically thick
 500 snowpacks (i.e., semi-infinite), because the errors of direct albedo in the visible are
 501 negligible compared with those in the near-IR. To improve the performance of 2-stream
 502 algorithms, we develop a parameterization that corrects the underestimated near-IR snow
 503 albedo at large zenith angles.

504

505 Figure 11 shows the direct near-IR albedo and fractional absorption of a 2-meter thick
 506 snowpacks consisting of grains with radius 100 μm and 1000 μm , computed using 2-
 507 stream algorithms and 16-stream DISORT. For solar zenith angles $> 75^\circ$, 2-stream
 508 models underestimate snow albedo and overestimate solar absorption within snowpack,
 509 mostly in the top 2-cm of snow. We define and compute R_{75+} as the ratio of direct semi-
 510 infinite near-IR albedo computed using 16-stream DISORT ($\alpha_{16-DISORT}$) to that computed
 511 using CICE (α_{CICE}). This ratio is shown in Figure 11 (c) and can be parameterized as a
 512 function of snow grain radius (r , unit in meter) and the cosine of incident solar zenith
 513 angle (μ_0), as shown in Figure 11(c):

514

$$515 \quad R_{75+} = \frac{\alpha_{16-DISORT}}{\alpha_{CICE}} = c_1(\mu_0) \log_{10}(r) + c_0(\mu_0) \quad (12)$$

516

517 where coefficients c_1 and c_0 are polynomial functions of μ_0 , as shown in Figure 11(d):

518

$$519 \quad c_1(\mu_0) = 1.304\mu_0^2 - 0.631\mu_0 + 0.086 \quad (13a)$$

$$520 \quad c_0(\mu_0) = 6.807\mu_0^2 - 3.338\mu_0 + 1.467 \quad (13b)$$

521

522 Since 2-stream models always underestimate snow albedo, R_{75+} always exceeds 1 (Figure
 523 11c). We can then adjust the direct near-IR snow albedo (α_{CICE}) and direct near-IR solar
 524 absorption ($Fabs_{CICE}$) by snow computed using CICE with ratio R_{75+} :

525

$$526 \quad \alpha_{CICE}^{adjust} = R_{75+} \alpha_{CICE} \quad (14a)$$

527

$$528 \quad Fabs_{CICE}^{adjust} = Fabs_{CICE} - (R_{75+} - 1) * \alpha_{CICE} * F_{nir} \quad (14b)$$

529



530 where F_{nir} is the direct near-IR flux. This adjustment reduces the error of near-IR albedo
531 from negative 2-10% to within $\pm 0.5\%$ for solar zenith angles larger than 75° , and for
532 grain radii ranging from 30-1500 μm (Figure 12). Errors in broadband direct albedo are
533 therefore also reduced to < 0.01 . The direct near-IR flux absorbed by the snowpack
534 decreases after applying this adjustment. In practice, if snow is divided into multiple
535 layers, we assume all decreased near-IR absorption (2nd term on the right hand side,
536 equation 14b) is confined within the top layer. This assumption is fairly accurate for the
537 near-IR band, since most direct IR absorption occurs at the very surface of snowpack
538 (Figures 10 and 11).

539

540 It is important to note that although the errors of direct near-IR albedos are large for large
541 solar zenith angles, the absolute error in reflected shortwave flux is small (Figures 7 and
542 8) as the down-welling solar flux reaches snowpack decreases as solar zenith angle
543 increases (Figures 1(b)). However, such small biases in flux can be important at high
544 latitudes where the solar zenith angle remains large for many days in late winter and early
545 spring. We have implemented this parameterization in MPAS-seaice to quantify its
546 impact on polar climate, though these experiments are beyond the scope of the present
547 paper.

548

549 **7. Implementation of snow radiative transfer model in Earth system models**

550

551 ESMs often use broader band-averaged SSPs of snow and aerosols for computational
552 efficiency, rather than using brute-force integration of spectral solar properties across
553 narrower bands (per equation 11). Besides using different radiative transfer
554 approximations, SNICAR and CICE also adopt different methods to derive the band-
555 averaged SSPs of snow for different band schemes.

556

557 In SNICAR, snow solar properties are computed for 5 bands: one visible band (0.3 -
558 0.7 μm), and four near-IR bands (0.7 - 1 μm , 1 - 1.2 μm , 1.2 - 1.5 μm , and 1.5 - 5 μm).
559 The solar properties of four subdivided near-IR bands are combined by fixed ratios to
560 compute the direct/diffuse near-IR snow properties. These two sets of ratios are derived
561 offline based on the incident solar spectra of typical of mid-latitude winter for clear and
562 cloudy-sky conditions clear sky and cloudy sky, respectively (Figure 1(a)).

563

564 The band-averaged SSPs of snow grains are computed following the Chandrasekhar
565 Mean approach (Thomas and Stamnes, 1999, their Equation 9.27; Flanner et al., 2007).



566 Specifically, spectral SSPs of snow grains are weighted into bands according to surface
 567 incident solar flux typical of mid-latitude winter for clear and cloudy sky conditions. In
 568 addition, the single-scattering albedo $\varpi(\lambda)$ of ice grains are also weighted by the
 569 hemispheric albedo $\alpha(\lambda)$ of an optically thick snowpack:

570

$$571 \quad \bar{\varpi}(\bar{\lambda}) = \frac{\int_{\lambda_1}^{\lambda_2} \varpi(\lambda) F(\lambda) \alpha(\lambda) d\lambda}{\int_{\lambda_1}^{\lambda_2} F(\lambda) \alpha(\lambda) d\lambda} \quad (15a)$$

$$572 \quad \bar{g}(\bar{\lambda}) = \frac{\int_{\lambda_1}^{\lambda_2} g(\lambda) F(\lambda) d\lambda}{\int_{\lambda_1}^{\lambda_2} F(\lambda) \alpha(\lambda) d\lambda} \quad (15b)$$

$$573 \quad \bar{\sigma}_{ext}(\bar{\lambda}) = \frac{\int_{\lambda_1}^{\lambda_2} \sigma_{ext}(\lambda) F(\lambda) d\lambda}{\int_{\lambda_1}^{\lambda_2} F(\lambda) \alpha(\lambda) d\lambda} \quad (15c)$$

574

575 Two sets of snow band-averaged SSPs are generated for all grain radii, suitable for direct
 576 and diffuse light, respectively. For each modeling step and band, SNICAR is called twice
 577 to compute the direct and diffuse snow solar properties.

578

579 In CICE, the snow-covered sea ice properties are computed for 3 bands: one visible band
 580 (0.3 – 07 μm), and two near-IR bands (0.7 – 1.19 μm and 1.19 – 5 μm). The solar
 581 properties of these two near-IR bands are combined using ratios w_{nir1} and w_{nir2} for 0.7-1
 582 .19 μm and 1.19-5 μm , depending on the fraction of direct near-IR flux f_{nidr} :

583

$$584 \quad w_{nir1} = 0.67 + 0.11 * (1 - f_{nidr}) \quad (16a)$$

$$585 \quad w_{nir2} = 1 - w_{nir1} \quad (16b)$$

586

587 The band SSPs of snow are derived by integrating the spectral SSPs and the spectral
 588 surface solar irradiance measured in the Arctic under mostly clear sky.

589

$$590 \quad \bar{\varpi}(\bar{\lambda}) = \int_{\lambda_1}^{\lambda_2} \varpi(\lambda) F(\lambda) d\lambda \quad (17a)$$

$$591 \quad \bar{g}(\bar{\lambda}) = \int_{\lambda_1}^{\lambda_2} g(\lambda) F(\lambda) d\lambda \quad (17b)$$

$$592 \quad \bar{\sigma}_{ext}(\bar{\lambda}) = \int_{\lambda_1}^{\lambda_2} \sigma_{ext}(\lambda) F(\lambda) d\lambda \quad (17c)$$

593



594 In addition, the band-averaged single-scattering albedo $\varpi(\bar{\lambda})$ is also increased to $\varpi(\bar{\lambda})'$
595 until the band albedo computed using averaged SSPs matches the band albedo $\bar{\alpha}$ within
596 0.0001, where $\bar{\alpha}$ is:

597

$$598 \quad \bar{\alpha} = \int_{\lambda_1}^{\lambda_2} \alpha(\lambda) F(\lambda) d\lambda \quad (18)$$

599

600 CICE adopts this single set of band SSPs for both direct and diffuse computations. In
601 practice, the physical snow grain radius r is adjusted to a radiatively equivalent radius r_{eqv}
602 based on the fraction of direct flux in the near-IR band (f_{nidr}):

603

$$604 \quad r_{eqv} = (f_{nidr} + 0.8(1 - f_{nidr}))r \quad (19)$$

605

606 This r_{eqv} and the corresponding snow SSPs are then used in the radiative transfer
607 calculation. The computed direct and diffuse solar properties alone are less accurate,
608 while the combined all-sky broadband solar properties agree with SNICAR (Briegleb and
609 Light, 2007). As a result, for each modeling step and band, CICE radiative transfer
610 subroutine is called only once to compute both the direct and diffuse snow solar
611 properties simultaneously.

612

613 SNICAR and CICE also use different approaches to avoid numerical singularities. In
614 SNICAR, singularities occur when the denominator of term C_n^\pm in equation (3) equals to
615 zero (i.e., $\gamma^2 - 1/\mu_0^2 = 0$), where γ is determined by the approximation method and SSPs
616 of snow, and μ_0 is the cosine of the solar zenith angle (Equations 23 and 24, Toon et al.,
617 1989). When such a singularity is detected, SNICAR will shift μ_0 by + 0.02 or -0.02 to
618 obtain physically realistic radiative properties. In the CICE algorithm, singularities arise
619 only when $\mu_0 = 0$ (Equation 4). Therefore, in practice, for $\mu_0 < 0.01$, CICE computes the
620 sea-ice solar properties for $\mu_0 = 0.01$ to avoid unphysical results.

621

622 **8. Discussion: a unified radiative transfer model for snow, sea ice, and land ice.**

623

624

625 Based on the inter-comparison of three 2-stream algorithms and their implementations in
626 ESMs, we formulated the following surface shortwave radiative transfer



627 recommendations for an accurate, fast, and consistent treatment for snow on land, land
628 ice, and sea ice in ESMS:

629

630 First, the 2-stream delta-Eddington adding-doubling algorithm by Briegleb and Light
631 (2007) is unsurpassed as a radiative transfer core. The evaluation in Section 5 shows that
632 this algorithm produces the least error for snow albedo and solar absorption within
633 snowpack, especially under overcast sky. This algorithm applies well to both uniformly
634 refractive media such as snow on land, and to non-uniformly refractive media, such as
635 bare/snow-covered/ponded sea ice and bare/snow-covered land ice. Numerical
636 singularities occur only rarely (when $\mu_0 = 0$) and are easily avoided in model
637 implementations. Among the three 2-stream algorithms discussed here, the CICE
638 radiative core is also the most efficient one as it takes only $\sim 2/3$ of the time of SNICAR
639 and 2SD to compute solar properties of multi-layer snowpacks.

640

641 Second, any 2-stream cryospheric radiative transfer model can incorporate the
642 parameterization described in Section 6 to adjust the low bias of direct near-IR snow
643 albedo and high bias of direct near-IR solar absorption in snow, for solar zenith angles
644 larger than 75° . These biases are persistent across all 2-stream algorithms discussed in
645 this work, and should be corrected for snow-covered surfaces. Alternatively, adopting a
646 4-stream approximation would reduce or eliminate such biases, though at considerable
647 expense in computational efficiency.

648

649 Third, a cryospheric radiative transfer model should prefer physically based
650 parameterizations that are extensible and convergent (e.g., with increasing spectral
651 resolution) for the band-averaged SSPs and size distribution of snow. Although the
652 treatments used in SNICAR and CICE are both practical since they both reproduce the
653 narrowband solar properties with carefully derived band-averaged inputs as discussed in
654 Section 7, the snow treatment used in SNICAR is more physically based and reproducible
655 since it does not rely on subjective adjustment and empirical coefficients as used in
656 CICE. Specifically, the empirical adjustment to snow grain radius implemented in CICE
657 may not always produce compensating errors. For example, in snow containing light-
658 absorbing impurities such adjustment may also lead to biases in aerosol absorption since
659 the albedo reduction caused by light-absorbing particles does not linearly depend on
660 snow grain radius (Dang et al., 2015). For further model development incorporating non-
661 spherical snow grain shapes (Dang et al., 2016; He et al., 2018ab), such adjustment on
662 grain radius may fail as well. Moreover, SNICAR computes the snow properties for four



663 near-IR bands, which helps capture the spectral variation of albedo (Figure 2) and
664 therefore better represents near-IR solar properties. It is also worth noting that unlike the
665 radiative core of CICE, SNICAR is actively maintained with numerous modifications and
666 updates in the past decade (e.g. Flanner et al., 2012; He et al., 2018b). Snow radiative
667 treatments that follow SNICAR conventions for SSPs may take advantage of these
668 updates. Note that any radiative core that follows SNICAR SSP conventions must be
669 called twice to compute diffuse and direct solar properties, respectively.

670

671 Fourth, a surface cryospheric radiative transfer model should flexibly accommodate
672 coupled simulations with distinct atmospheric and surface spectral grids. Both the 5-band
673 scheme used in SNICAR and the 3-band scheme used in CICE separate the visible from
674 near-IR spectrum at 0.7 μm . This boundary aligns with the Community Atmospheric
675 Model's original radiation bands (CAM; Neale et al., 2012), though not with the widely
676 used Rapid Radiative Transfer Model (RRTMG; Iacono et al., 2008) which places 0.7 μm
677 squarely in the middle of a spectral band. A mismatch in spectral boundaries between
678 atmospheric and surface radiative transfer schemes can require an ESM to unphysically
679 apportion energy from the straddled spectral bin when coupling fluxes between surface
680 and atmosphere. The spectral grids of surface and atmosphere radiation need not be
681 identical so long as the coarser grid shares spectral boundaries with the finer grid. In
682 practice maintaining a portable cryospheric radiative module such as SNICAR requires a
683 complex offline toolchain (Mie solver, spectral refractive indices for air, water, ice, and
684 aerosols, spectral solar insolation for clear and cloudy skies) to compute, integrate, and
685 rebin SSPs. Aligned spectral boundaries between surface and atmospheric would simplify
686 the development of efficient and accurate radiative transfer for the coupled Earth system.

687

688 Last, it is important to note that, although we only examine the performance of the CICE
689 adding-doubling algorithm for pure snow in this work, this algorithm can be applied to
690 the surface solar calculation of all cryospheric components with or without light-
691 absorbing particles present. First, Briegleb and Light (2007) proved its accuracy for
692 simulating ponded/bare sea-ice solar properties against observations and a Monte Carlo
693 radiation model. Second, In CESM and E3SM, the radiative transfer simulation of snow
694 on land ice is carried out by SNICAR with prescribed land ice albedo. Adopting the
695 CICE adding-doubling core in SNICAR will permit these ESMs to couple the snow and
696 land ice as a non-uniformly refractive column for more accurate solar computations since
697 bare/snow-covered/ponded land ice is physically similar to bare/snow-covered/ponded
698 sea ice, and the latter is already treated well by CICE radiative transfer core. Third,



699 adding light-absorbing particles in snow will not change our results qualitatively. Both
700 CICE and SNICAR simulate the impact of light-absorbing particles (black carbon and
701 dust) on snow and/or sea ice using self-consistent particle SSPs that follow the SNICAR
702 convention. The adoption of CICE radiative transfer algorithm in SNICAR, and the
703 implementation of SNICAR snow SSPs in CICE will enable a consistent simulation on
704 the radiative effects of light-absorbing particles in the cryosphere across ESM
705 components.

706

707 In summary, this inter-comparison and evaluation has shown multiple ways that the solar
708 properties of cryospheric surfaces can be improved in the current generation of ESMs.
709 We have adopted these recommendations in a hybrid model SNICAR-AD, implemented
710 in MPAS-seaice and E3SM Land Model (ELM), to examine the response of climate to
711 this improved and unified cryospheric surface radiation treatment in future E3SM studies.

712

713 **9. Conclusions**

714

715 In this work, we aim to improve and unify the solar radiative transfer calculations for
716 snow on land and snow on sea ice in ESMs by evaluating the following 2-stream
717 radiative transfer algorithms: the 2-stream delta-Eddington adding-doubling algorithm
718 implemented in sea-ice model Icepak/CICE/MPAS-seaice, the 2-stream delta-Eddington
719 and 2-stream delta-Hemispheric-Mean algorithms implemented in snow model SNICAR,
720 and a 2-stream delta-Discrete-Ordinate algorithm. Among these three models, the 2-
721 stream delta-Eddington adding-doubling algorithm produces the most accurate snow
722 albedo and solar absorption (Section 5). All 2-stream models underestimate near-IR snow
723 albedo and overestimate near-IR absorption when solar zenith angles are larger than 75°,
724 which can be adjusted by a parameterization we developed (Section 6). We compared the
725 implementations of radiative transfer cores in SNICAR and CICE (Section 7) and
726 recommended a consistent shortwave radiative treatment for snow-covered surfaces
727 across ESMs (Section 8). Improved treatment of surface cryospheric radiative properties
728 in the thermal infrared has recently been shown to remediate significant climate
729 simulation biases in Polar Regions (Huang et al., 2018). It is hoped that adoption of
730 improved and consistent treatments of solar radiative properties for snow-covered
731 surfaces as described in this study (i.e. the hybrid model SNICAR-AD) will further
732 remediate simulation biases in Polar Regions.

733



734 **Data availability.** The data and models are available upon request to Cheng Dang
735 (cdang5@uci.edu). SNICAR and CICE radiative transfer core can be found at
736 <https://github.com/E3SM-Project/E3SM>.
737

738 **Competing interests.** The authors declare that they have no conflict of interest.
739

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746

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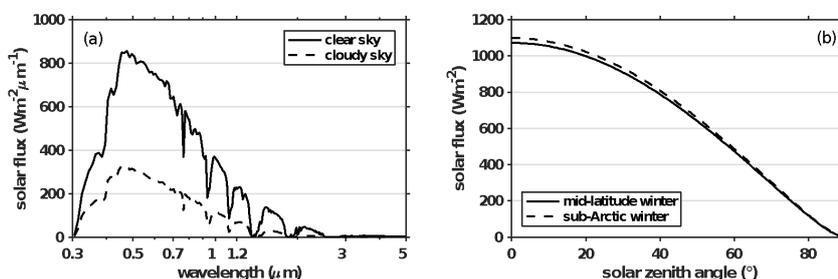
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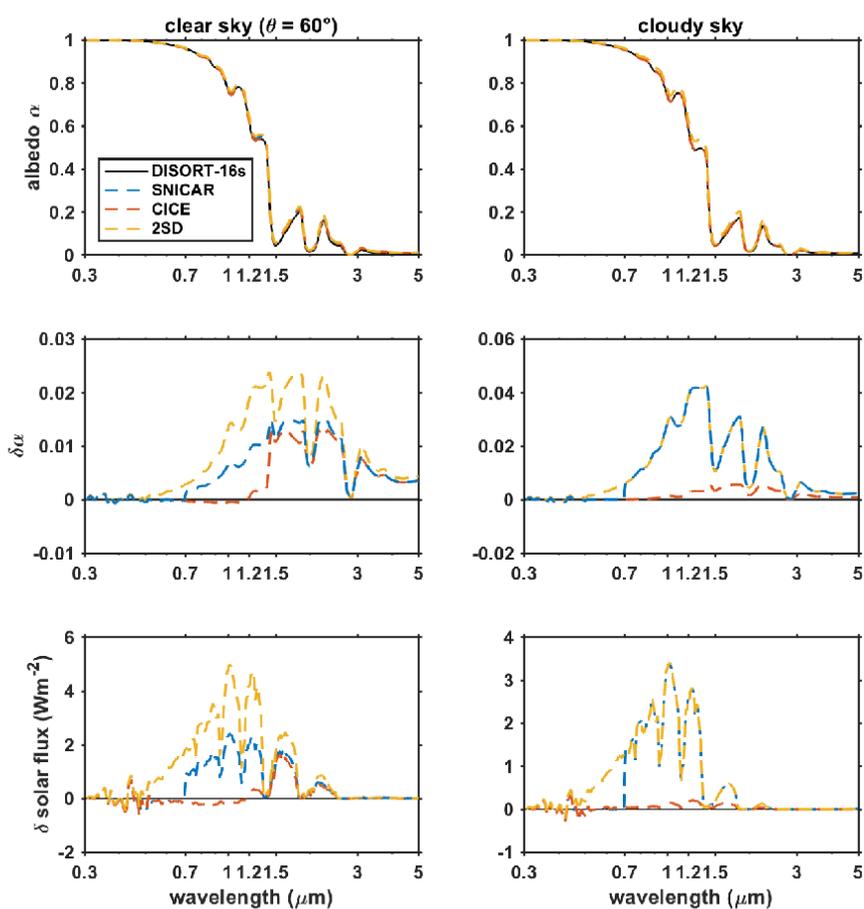
914 Figure 1. Spectral and total down-welling solar flux at surface computed using SWNB2
915 for (a) standard clear-sky and cloudy-sky atmospheric profiles of mid-latitude winter
916 assuming solar zenith angle is 60° at the top of atmosphere, and for (b) standard clear sky
917 profiles of mid-latitude and sub-Arctic winter with different incident solar zenith angles.
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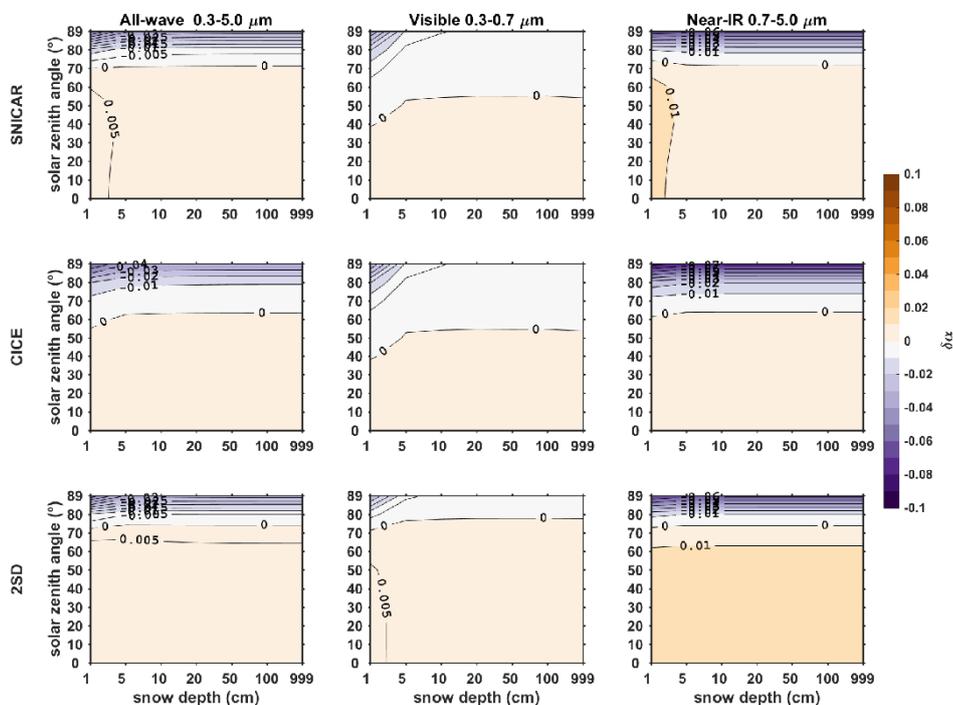
924 Figure 2. Spectral albedo of pure snow computed using 16-stream DISORT, SNICAR,
 925 CICE, and 2SD models, for clear-sky (direct beam at solar zenith angle 60°) and cloudy-
 926 sky conditions in the left and right panels, respectively. The top panels show spectral
 927 albedo. The middle panels show the difference ($\delta\alpha = \alpha_s - \alpha_{16}$) in spectral albedos
 928 computed using 2-stream model (α_s) and 16-stream DISORT (α_{16}). The bottom panels
 929 show the different of reflected spectral flux given $\delta\alpha$. The snowpack is set to semi-
 930 infinite deep with grain radius of $100 \mu\text{m}$.
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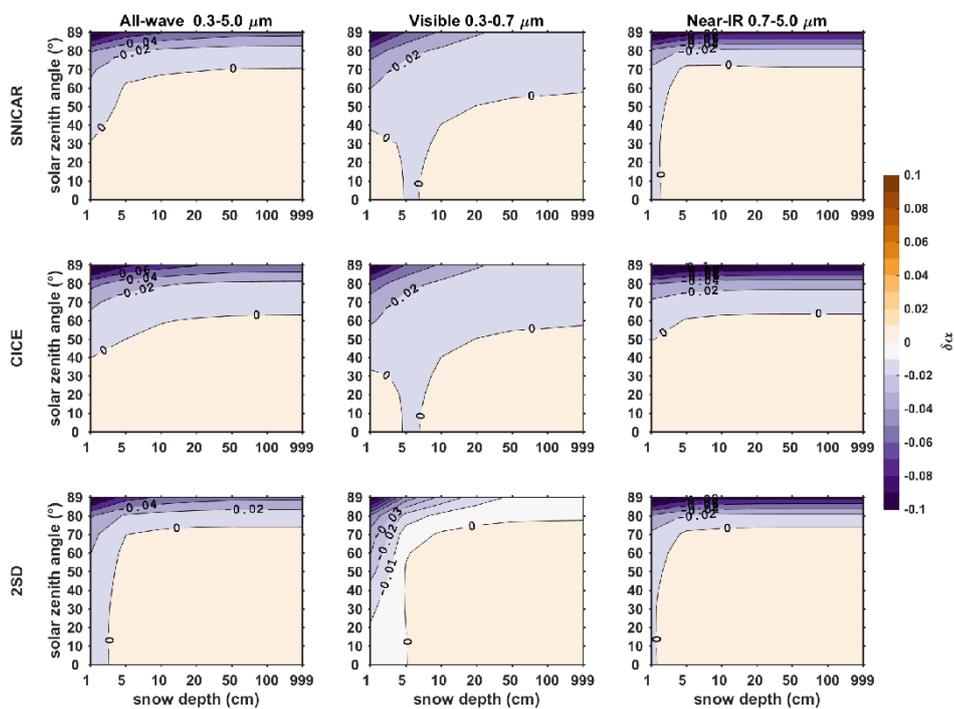
935 Figure 3. The difference in direct snow albedo ($\delta\alpha = \alpha_2 - \alpha_{16}$) computed using 2-stream
 936 models (α_2) and using 16-stream DISORT model (α_{16}), for various snow depths and solar
 937 zenith angles, with snow gran radius of 100 μm . From the top to the bottom rows are
 938 results of 2-stream models SNICAR, CICE, and 2SD. From the left to the right columns
 939 are albedo differences of all-wave, visible, near-IR bands.



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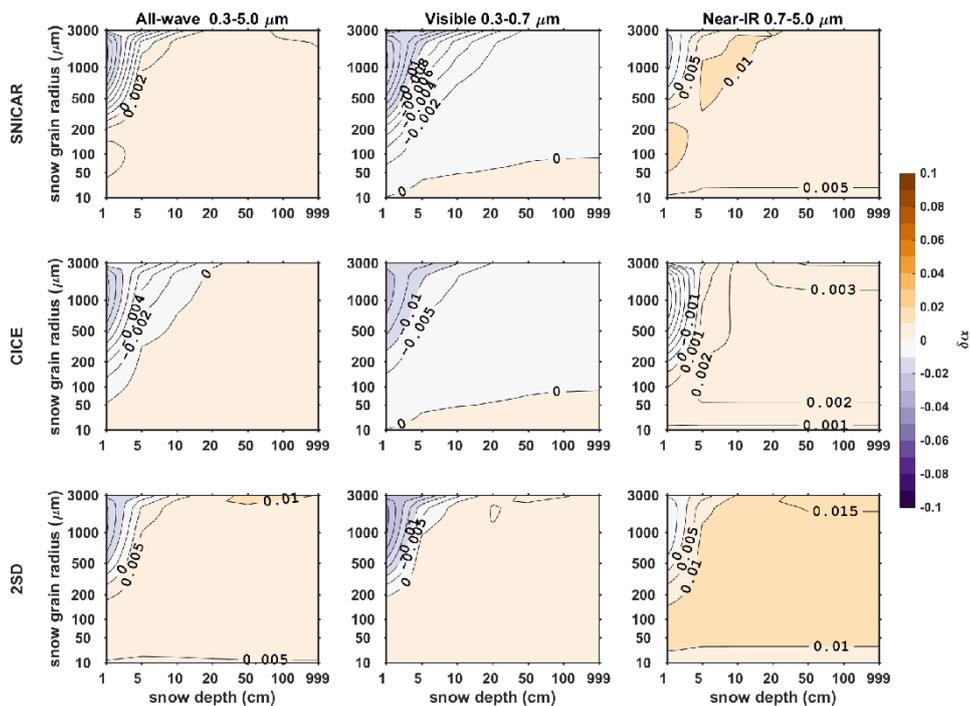
942 Figure 4. The same to Figure 3, but for snow grain radius of 1000 μm .
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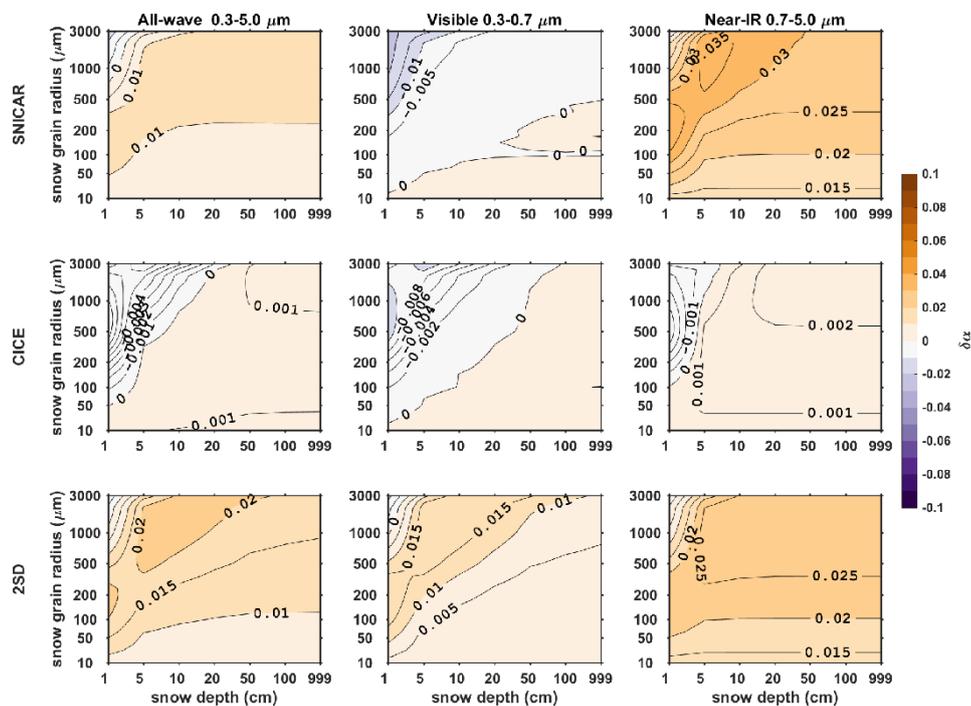
947 Figure 5. The same to Figure 3, but for fixed solar zenith angle of 60° and different snow
948 grain radii.
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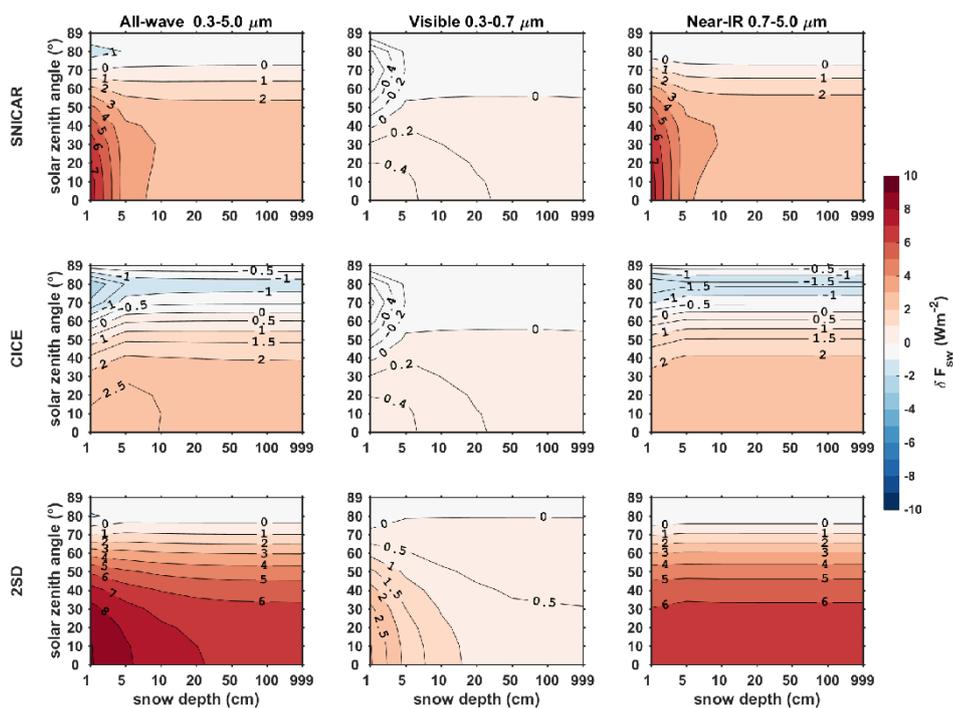
952 Figure 6. The same to Figure 5, but for diffuse snow albedo with different snow grain
953 radii.
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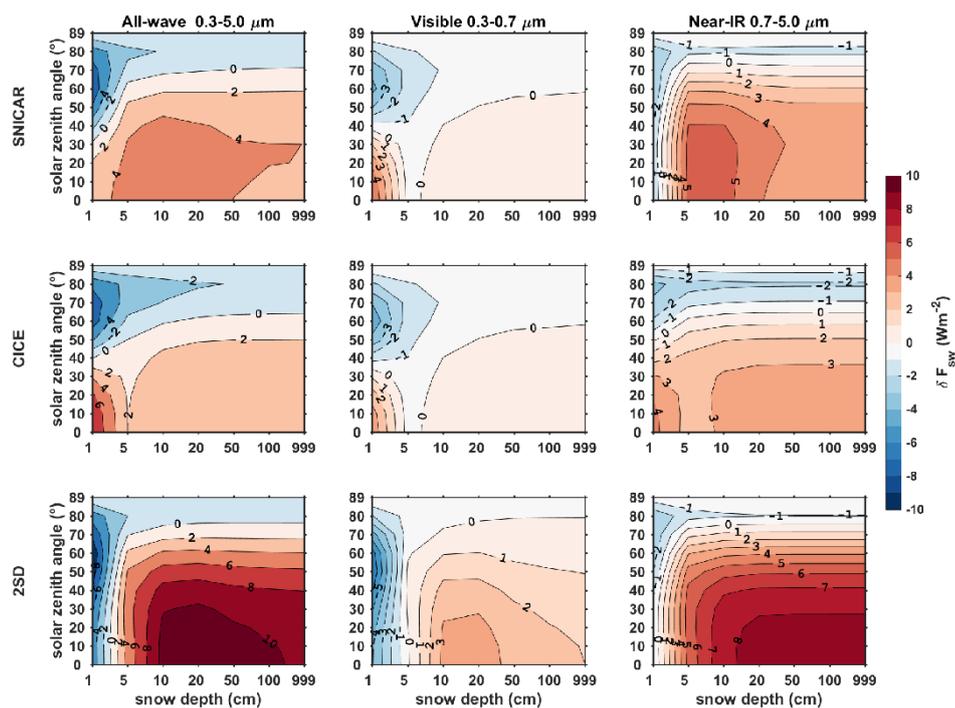
957 Figure 7. Error in reflected direct solar flux given albedo errors shown in Figure 3.



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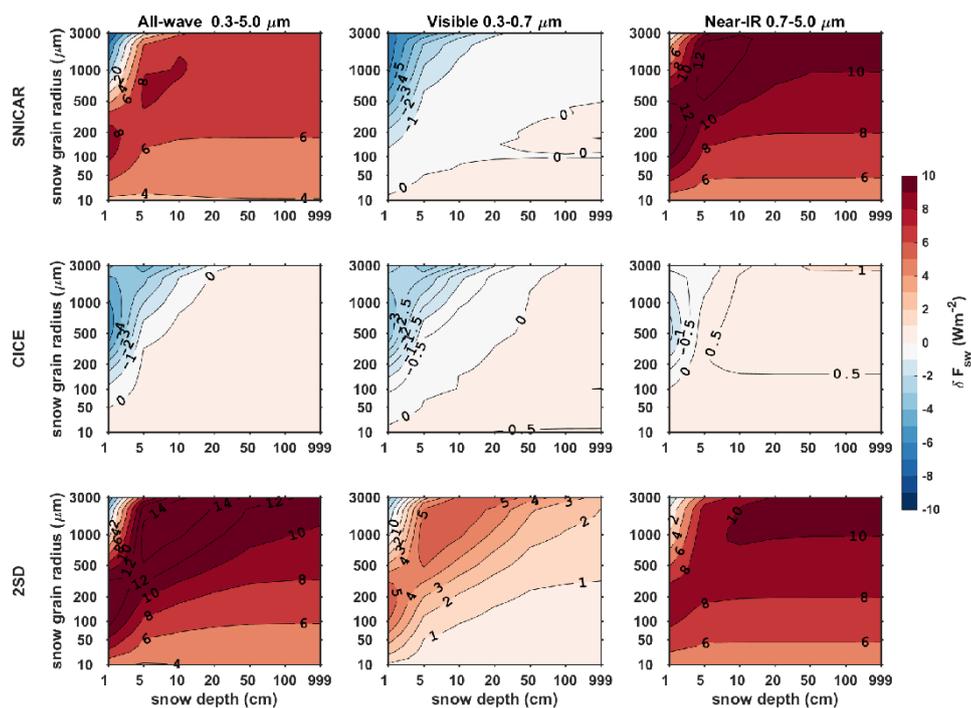
960 Figure 8. Error in reflected direct solar flux given albedo errors shown in Figure 4.



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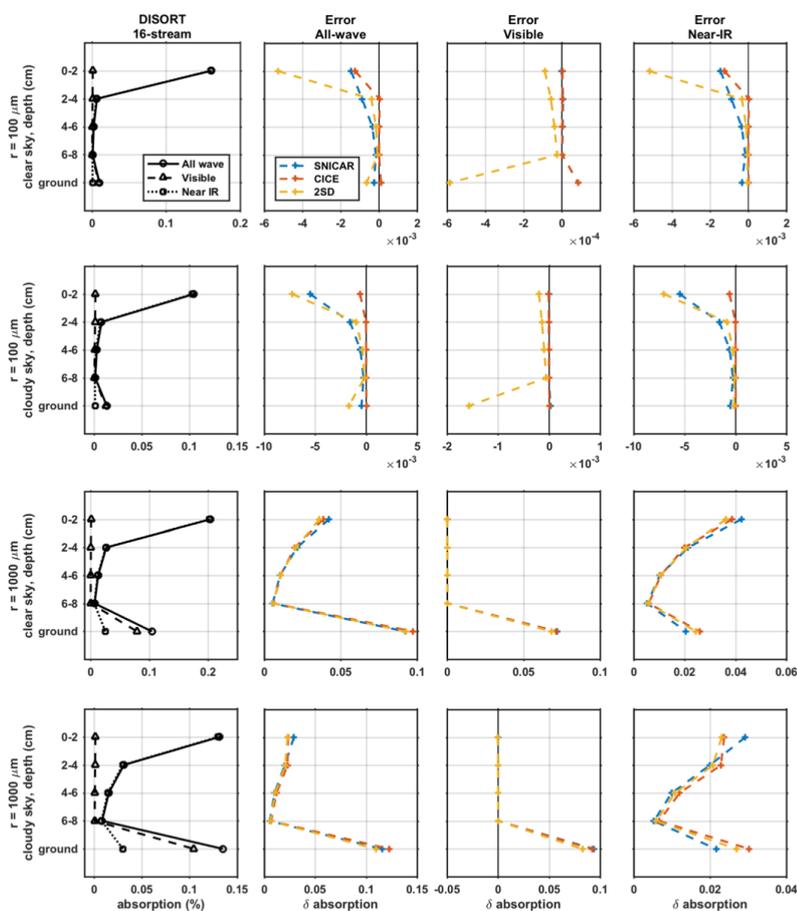
963 Figure 9. Error in reflected diffuse solar flux given albedo errors shown in Figure 6.



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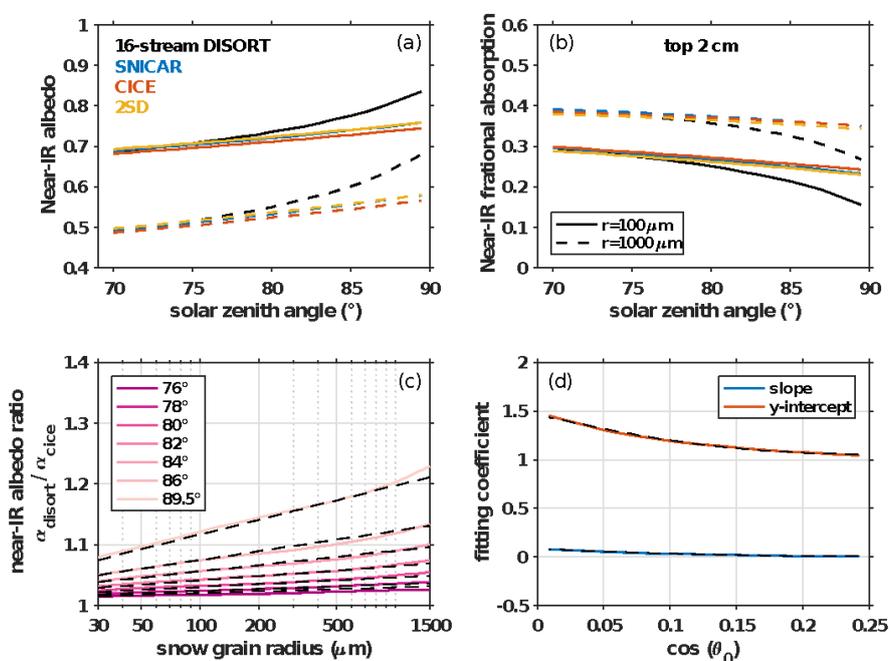
967 Figure 10. Comparison of light-absorption profiles derived from 2-stream models and 16-
 968 stream DISORT. The left-most column show fractional band absorptions computed using
 969 16-stream DISORT. The right three panels show the errors of all-wave, visible, and near-
 970 IR fractional absorptions calculated using 2-stream models. The top and bottom panels
 971 are for clear-sky and cloudy-sky conditions (solar zenith angle of 60°), respectively. The
 972 snowpack is 10 cm deep, and is divided evenly into five 2-cm thick layers, for new snow
 973 ($r = 100 \mu\text{m}$) and old snow ($r = 1000 \mu\text{m}$). The layers 1-4 represent the top four snow
 974 layers (top 8 cm), and layer 5 represents underlying ground with albedo of 0.25.
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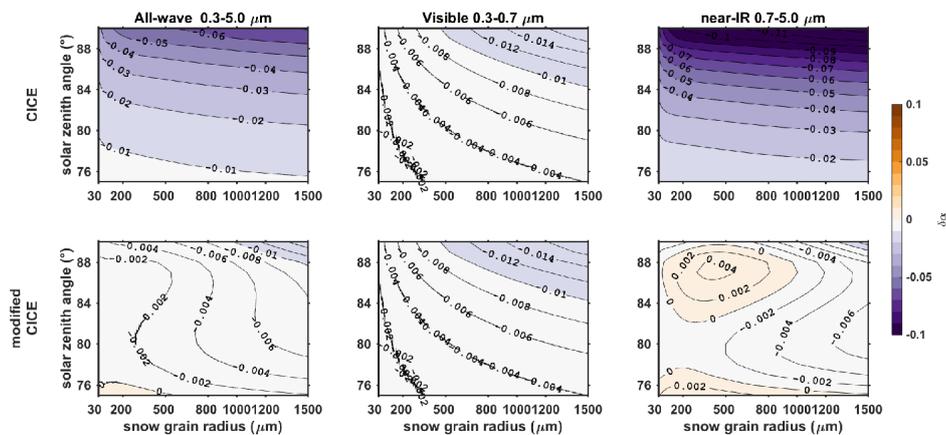
979 Figure 11. (a) Direct near-IR snow albedo and (b) near-IR fractional absorption by top 2-
 980 cm snow of a 2-m thick snowpack, for solar zenith angles larger than 70° and snow grain
 981 radii of $100\ \mu\text{m}$ and $1000\ \mu\text{m}$. (c) The ratios of near-IR albedo computed using CICE to
 982 that computed using 16-stream DISORT for different solar zenith angles. These ratios are
 983 parameterized as liner functions of the logarithmic of snow grain radius. The slopes and
 984 y-intercepts are shown in (d). The black dashed curves in figures (c) and (d) are fitting
 985 values computed using parameterization discussed in Section 5.
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989 Figure 12. Error in semi-infinite snow albedo computed using CICE before (top row) and
 990 after (bottom row) incorporating corrections for near-IR albedo, for different solar zenith
 991 angles and snow grain radii.
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996 Table 1. Two-stream radiative transfer algorithms evaluated in this work, including
 997 algorithms that are currently implemented in Earth System Model CESM and E3SM.
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ESM Component	Land	Sea Ice	
Model	SNICAR	CICE/MPAS-seaice	2SD
Radiative transfer approximation	2-stream δ -Eddington (visible) δ -Hemispheric-mean (near-IR)	2-stream δ -Eddington	2-stream δ -Discrete-ordinate
Treatment for multi-layered media	matrix inversion	adding-doubling	matrix inversion
Fresnel reflection/refraction	no	yes	yes
Number of bands implemented in ESMs	5 bands (1 visible, 4 near-IR)	3 bands (1 visible, 2 near-IR)	
Applies to	snow	bare/ponded/snow-covered sea ice, and snow	bare/ponded/snow-covered sea ice, and snow

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